## Comment on 'The effect of rotation on the Rayleigh–Bénard stability threshold' by A. Prosperetti *Phys. Fluids* **24**,114101 (2012)

J.H.P. Dawes

Department of Mathematical Sciences, University of Bath, Claverton Down, Bath BA2 7AY, UK

In this Comment we discuss the results presented in Refs. [5] and [4] for the classic problem of Rayleigh– Bénard convection, i.e. the thermal instability of a horizontal layer of viscous fluid heated uniformly from below. In the case of fixed temperature and stress-free horizontal boundaries the critical dimensionless temperature difference (Rayleigh number R) for the onset of motion can be determined exactly ([6]). For fixed temperature, but rigid, boundaries the computation becomes more involved. This case was treated by Chandrasekhar, see Chapter II, section 17 of [1]. He formulated the stability problem as a variational principle and obtained an approximate solution through the use of an appropriate solution *ansatz*. This leads to the following approximate expression for (an upper bound to) the critical Rayleigh number, his equation (311):

$$R(k) = \frac{\beta^6}{k^2} \left[ 1 - \frac{16\pi^2 k \cosh^2(k/2)}{\beta^4 (k + \sinh k)} \right]^{-1},\tag{1}$$

where  $\beta^2 \equiv \pi^2 + k^2$  and k is the horizontal wavenumber of the perturbation to the conduction solution. This expression was re-derived in Ref. [4] through a similar approximation procedure, but without being formulated as a variational problem, and without reference to its earlier existence in Chandrasekhar's book.

We now turn to the effect of rotation on the Rayleigh-Bénard instability for rigid, fixed temperature boundaries. Chandrasekhar considers the effect of rotation about a vertical axis in Chapter III of [1]. This leads naturally to a very similar instability problem, with the overall effect of rotation being to stabilise the flow. Although well-studied, interest in obtaining simplified descriptions of the linear instability continues to attract attention, see for example [3]. We denote the dimensionless rotation rate (Taylor number) by T. The linear instability of thermal convection with rotation can also be formulated as a variational problem, leading again to an approximation for the marginal stability curve R(k,T). As in the non-rotating case, the variational formulation guarantees that any estimate of the critical Rayleigh number will be an upper bound on the true value. Chandrasekhar does not present as simple an explicit expression as (1) but implicitly it is his equation (159), and it is evaluated numerically in his Table VIII. The one-term approximation he presents can be written in the reasonably compact form

$$R(k,T) = \frac{\beta^6 + \pi^2 T}{k^2} \left[ 1 + \frac{4\pi^2 F}{\beta^2 (\beta^6 + \pi^2 T)} \right]^{-1},$$
(2)

where

$$F = \frac{\sum_{\text{cyclic}} q_1 \beta_1^4 (q_3^2 - q_2^2) (q_2^2 - k^2) (q_3^2 - k^2) \tanh(q_1/2)}{\sum_{\text{cyclic}} (q_1^2 - k^2) q_2 q_3 (q_3^2 - q_2^2) \tanh(q_2/2) \tanh(q_3/2)},$$
(3)

is the part of the expression that indicates the effect of the rigid boundaries (note that the exact result for the stress-free case is recovered if F = 0). In (3) the sums are taken over the three terms obtained by cyclically permuting the subscripts (1 2 3), and we define  $\beta_j = \pi^2 + q_j^2$ . The coefficients  $q_1, q_2, q_3$  are defined as follows:  $q_1^2, q_2^2$  and  $q_3^2$  are the distinct roots of the cubic polynomial  $(q^2 - k^2)^3 + Tq^2 = 0$ .

defined as follows:  $q_1^2$ ,  $q_2^2$  and  $q_3^2$  are the distinct roots of the cubic polynomial  $(q^2 - k^2)^3 + Tq^2 = 0$ . The need to compute the roots of the cubic is an additional complication, Using standard methods, and defining  $Y = (12\sqrt{12T^3 + 81T^2k^4} - 108Tk^2)^{1/3}$ , we obtain, exactly,

$$q_1^2 = k^2 + Y/6 - 2T/Y, \qquad q_2^2 = k^2 - e^{i\pi/3}(Y/6 + 2T/Y), \qquad q_3^2 = k^2 - e^{-i\pi/3}(Y/6 + 2T/Y).$$
 (4)

Alternatively, one can compute very good explicit approximations to the roots in a number of ways, for example through applying a Shanks transformation to the first three terms of the Taylor series solution, using  $T^{-1}$  as the small parameter [2]. The Shanks transformation yields the simpler, but approximate, expressions

$$q_1^2 = \frac{k^6}{T+3k^4}, \qquad q_2^2 = k^2 + \frac{5k^2T^{1/2} + 4iT}{4T^{1/2} - 3ik^2}, \qquad q_3^2 = k^2 + \frac{5k^2T^{1/2} - 4iT}{4T^{1/2} + 3ik^2}.$$
 (5)

In Ref. [5] the alternative expression

$$R = \left(\frac{\beta^6}{k^2} + (1-b)T\right) \left[1 - \frac{8\pi^2 k (1+\cosh k)}{\beta^4 (k+\sinh k)}\right]^{-1},\tag{6}$$

is proposed, see equation (25) in that paper, and where

$$b = k \left( \frac{8\pi^2}{\beta^4} \frac{1 + \cosh k}{k + \sinh k} + \frac{k - \sinh k}{(k + \tanh k)k^2 \tanh(k/2)} - \frac{1}{k + \sinh k} + \frac{4}{\beta^2 \tanh(k/2)} \right),\tag{7}$$

i.e. in the notation of [5],  $b = \beta^6 \tilde{B}_{11}/k^2$  where  $\tilde{B}_{11}$  is defined in [5], Appendix A, equation (A23). The expressions (6) - (7) are not as complicated as (2) - (3), but they are also not as accurate. This is demonstrated clearly in the tables below where we compare the critical Rayleigh numbers  $R_c$  and wavenumbers  $k_c$  over a range of values of T.

Table 1: Comparison of estimates for the marginal stability threshold  $R_c(T)$  and the corresponding critical wavenumber  $k_c(T)$  for the onset of convection in the presence of rotation. Figures in **bold** differ from the (exact) numerical values quoted in Table I of [5].

	Numerical		(2)-(3)-(5)		(6)-(7)	
T	$R_c(T)$	$k_c(T)$	$R_c(T)$	$k_c(T)$	$R_c(T)$	$k_c(T)$
$10^{2}$	1 756.3	3.16	$1\ 763.8$	3.16	$1\ 753.3$	3.15
$10^{3}$	2  151.3	3.49	$2\ 159.3$	3.48	$2\ 074.7$	3.41
$10^{4}$	4 712.0	4.79	$4\ 717.1$	4.78	$4\ 420.6$	4.69
$10^{5}$	16  719	7.17	$16\ 743$	7.18	$17\ 100$	7.56
$10^{6}$	71  085	10.82	71 587	10.88	78  569	11.92
$10^{7}$	324 510	16.34	328  180	16.47	$371 \ 470$	18.16
$10^{8}$	$1 \ 525 \ 100$	24.64	$1 \ 543 \ 900$	24.83	$1\ 760\ 700$	27.20
$10^{9}$	$7\ 244\ 600$	37.01	$7 \ 326 \ 200$	37.27	$8 \ 316 \ 800$	40.40
$10^{10}$	$34 \ 498 \ 000$	<b>55.40</b>	$34 \ 817 \ 000$	55.71	$39\ 118\ 000$	59.74

Table 2: Comparison of percentage errors between numerical values and those given by different approximation expressions for  $R_c(T)$  and  $k_c(T)$ . Figures in **bold** indicate substantial differences from those quoted in Table I of [5].

	Numerical		(2)-(3)-(5)		(6)-(7)	
Т	$R_c(T)$	$k_c(T)$	$\% R_c$	$\% k_c$	$\% R_c$	$\% k_c$
$10^{2}$	$1\ 756.3$	3.16	0.42	0.93	-0.18	-0.41
$10^{3}$	$2\ 151.3$	3.49	0.37	0.06	-3.56	-2.12
$10^{4}$	$4\ 712.0$	4.79	0.11	0.02	-6.19	-1.90
$10^{5}$	$16\ 719$	7.17	0.14	0.20	2.28	5.47
$10^{6}$	71  085	10.82	0.70	0.54	10.53	10.21
$10^{7}$	324  510	16.34	1.13	0.76	14.47	11.14
$10^{8}$	$1 \ 525 \ 100$	24.64	1.24	0.79	15.45	10.42
$10^{9}$	$7\ 244\ 600$	37.01	1.13	0.69	14.80	9.17
$10^{10}$	$34 \ 498 \ 000$	55.40	0.93	0.55	13.39	7.83

Table 1 shows the values for the critical Rayleigh number  $R_c$  at which R attains a minimum over wavenumbers k at fixed T, computed numerically, and estimated by the two different methods. Table 2 shows the percentage errors in both  $R_c$  and the corresponding critical wavenumber  $k_c$  for both approximate expressions. While those for (2) using (3) and (5) for the roots are around 1% accurate (or better) at all values of T shown, expression (6) is substantially worse, particularly at moderately high T. Qualitatively, we remark also that (i) the approximate values computed using the formula (6) of Ref [5] are neither consistently above nor consistently below the true values, and (ii) that the approximation appears to improve only relatively slowly at large T; as shown in [5] the approximation (6) yields the well-known asymptotic limits for  $R_c$  and  $k_c$  as  $T \to \infty$ , therefore both methods must become increasingly accurate in the limit  $T \to \infty$ . On the other hand, comparison of the marginal stability curves produced by the two methods confirms that (6) is more accurate at low Taylor numbers,  $T \leq 10^2$ , where the exact roots (4) must be used instead of (5) to preserve the accuracy of Chandrasekhar's approximation.

## References

- S. Chandrasekhar, Hydrodynamic and Hydromagnetic Stability (Clarendon, Oxford, 1961) [reprinted by Dover, New York, 1981].
- [2] E.J. Hinch, *Perturbation Methods* (Cambridge University Press, Cambridge, 1991).
- R.C. Kloosterziel and G.F. Carnevale, Closed-form linear stability conditions for rotating Rayleigh– Bénard convection with rigid stress-free upper and lower boundaries. J. Fluid Mech. 480, 25–42 (2003)
- [4] A. Prosperetti, A simple analytic approximation to the Rayleigh-Bénard stability threshold. *Phys. Fluids* 23, 124101 (2011)
- [5] A. Prosperetti, The effect of rotation on the Rayleigh-Bénard stability threshold. *Phys. Fluids* 24, 114101 (2012)
- [6] Lord Rayleigh, On convection currents in a horizontal layer of fluid when the higher temperature is on the under side. *Phil. Mag.* 32, 529–546 (1916)