

THE GEOMETRY OF THE TODA EQUATION

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ABSTRACT. I show that solutions of the $SU(\infty)$ Toda field equation generating a fixed Einstein-Weyl space are governed by a linear equation on the Einstein-Weyl space. From this, obstructions to the existence of Toda solutions generating a given Einstein-Weyl space are found. I also give a classification of Einstein-Weyl spaces arising from the Toda equation in more than one way. This classification coincides with a class of spaces found by Ward and hence clarifies some of their properties. I end by discussing the simplest examples.

1. INTRODUCTION

In [16], Ward showed that solutions $u(x, y, z)$ of the $SU(\infty)$ Toda field equation $u_{xx} + u_{yy} + (e^u)_{zz} = 0$ may be used to define three dimensional Einstein-Weyl spaces. A Weyl space is a conformal manifold M together with a compatible torsion-free connection (called a *Weyl connection*) and it is said to be *Einstein-Weyl* iff the symmetric tracefree part of the Ricci tensor of this connection vanishes [7]. Weyl connections on a conformal manifold correspond bijectively to covariant derivatives (called Weyl derivatives) on the density line bundle L^1 , which is the oriented real line bundle whose n th power is $|\Lambda^n TM|$ where $n = \dim M$.

A Weyl space may be described by a choice of compatible Riemannian metric g and the connection 1-form ω of the Weyl derivative on L^1 relative to the trivialisation of L^1 determined by the volume form of the metric (so that, for the induced Weyl connection, $Dg = -2\omega \otimes g$). In these terms, the Einstein-Weyl space defined by the solution u of the Toda equation may be written:

$$(1.1) \quad \begin{aligned} g &= e^u(dx^2 + dy^2) + dz^2 \\ \omega &= -u_z dz. \end{aligned}$$

The Toda equation is a nonlinear integrable system, but very few solutions are known explicitly [1, 3, 4, 13]. Ward found an implicit procedure for generating a family of solutions from axially symmetric harmonic functions V . The Einstein-Weyl spaces determined by these implicit solutions are nevertheless completely explicit (in terms of V) and Ward suggested that “. . . further investigation is needed to clarify the nature and properties of this family of spaces” [16].

In this paper I show that these Einstein-Weyl spaces are precisely the Einstein-Weyl spaces which can be written in the form (1.1) in at least two inequivalent

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ways. The key observation is that the solutions of the Toda equation on a fixed Einstein-Weyl background are essentially given by solutions of a linear system in this Weyl geometry. More precisely, the solutions of this linear system, the ‘‘Toda structures’’, correspond to solutions of the Toda equation on the Einstein-Weyl space up to changes of isothermal coordinates (x, y) and translation of z . As a consequence, I show that an Einstein-Weyl space admits at most a four dimensional space of compatible Toda structures, with equality iff the Einstein-Weyl space is Einstein. Furthermore, obstructions to the existence of Toda structures on a given Einstein-Weyl space are found. These obstructions are sufficient to establish which of the local forms of compact Einstein-Weyl spaces (found in [12]) admit Toda structures.

In section 3, I prove that the existence of more than one Toda structure on an Einstein-Weyl space is equivalent to the existence of a conformal vector field of a special type, which I will call an ‘‘axial symmetry’’. This fact is used in section 4 to classify the resulting spaces. The simplest examples are then discussed in the final section.

I work throughout with the density bundles L^w ($w \in \mathbb{R}$). A conformal structure may then be defined as an L^2 valued metric, so that the conformal inner product of vector fields X, Y is $\langle X, Y \rangle \in C^\infty(M, L^2)$. Compatible Riemannian metrics correspond to trivialisations of L^1 , and such a trivialisation is often called a *length scale* or *gauge*. When tensoring with a density line bundle, I shall omit the tensor product sign, and sections of $L^{w-1}TM$ or $L^{w+1}T^*M$ are called vector fields or 1-forms of weight w respectively. The Hodge star operator on an oriented conformal 3-manifold identifies L^w with $L^{w+3}\Lambda^3T^*M$ and $L^{w+1}T^*M$ with $L^{w+2}\Lambda^2T^*M$ and it will be taken to have square $-id$. For further details see [2, 5]. The results in this paper are local in character, and so, where necessary, vector fields are taken to be nonvanishing and manifolds simply connected.

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2. TODA STRUCTURES ON EINSTEIN-WEYL SPACES

The Einstein-Weyl spaces arising from the Toda Ansatz (1.1) have been characterised by Tod [13] as those which admit a shear-free twist-free geodesic congruence. If $\chi \in C^\infty(M, L^{-1}TM)$ denotes the weightless unit vector field tangent to this congruence (an oriented foliation with one dimensional leaves) then this means that

$$(2.1) \quad D\chi = \tau(id - \langle \chi, \cdot \rangle \otimes \chi)$$

where D is the Weyl connection, τ is a section of L^{-1} and $\langle \chi, \cdot \rangle$ denotes the weightless 1-form dual to χ with respect to the conformal structure. In (1.1), the congruence generated by $\partial/\partial z$ has this property and one finds that $2\tau\langle \chi, \cdot \rangle = -u_z dz$. Hence the Weyl derivative $D - 2\tau\langle \chi, \cdot \rangle$ is induced by the Levi-Civita connection of g and so the metric g is canonically determined, up to a constant multiple, by the Weyl structure and the congruence [2, 13]. I will denote by μ the trivialisation of L^1 corresponding to g and refer to this gauge μ (unique up to a constant)

as the *LeBrun-Ward* gauge [10, 16]. The congruence χ determines (in principle) the solution of the Toda equation up to the choice of isothermal coordinates (x, y) and affine changes of z . Fixing the LeBrun-Ward gauge μ determines z up to translation. In (1.1), $\tau = -\frac{1}{2}u_z\mu^{-1}$, $\chi = \mu^{-1}\partial/\partial z$ and $\langle \chi, \cdot \rangle = \mu dz$.

The equation (2.1) for χ is apparently nonlinear, but it actually becomes linear as an equation for the weight $\frac{1}{2}$ vector field $\mathcal{X} = \mu^{1/2}\chi$: explicitly, $D\mathcal{X} = \sigma id$, for some section $\sigma = \mu^{1/2}\tau$ of $L^{-1/2}$. Conversely, if $D\mathcal{X}$ is a multiple of the identity, then $\chi = \mathcal{X}/|\mathcal{X}|$ is a shear-free twist-free geodesic congruence and $\mu = |\mathcal{X}|^2$ is the LeBrun-Ward gauge.

Although this observation is trivial, it is the key idea behind the results of this paper, and so I should explain its origins. In [10], LeBrun gave a characterisation of the Toda Einstein-Weyl spaces in terms of *minitwistor theory* [7]. The space of oriented geodesics in a three dimensional Einstein-Weyl space is a complex surface \mathcal{S} containing rational curves (“minitwistor lines”) with normal bundle $\mathcal{O}(2)$, and shear-free geodesic congruences correspond to divisors in \mathcal{S} of degree 2 on each minitwistor line. LeBrun noticed that if the congruence is also twist-free, then the corresponding divisor is actually a divisor for $K_{\mathcal{S}}^{-1/2}$, where $K_{\mathcal{S}}$ is the canonical bundle of \mathcal{S} . After incorporating the choice of homothety factor of the LeBrun-Ward gauge, Toda structures on a fixed Einstein-Weyl space correspond to holomorphic sections of $K_{\mathcal{S}}^{-1/2}$. This immediately suggests that a linear equation is involved, and by applying the Penrose transform, following Tsai [15], one finds that sections of $K_{\mathcal{S}}^{-1/2}$ correspond to weight $\frac{1}{2}$ vector fields with tracelike derivative. It is then not hard to guess the relationship between such a vector field and χ .

2.1. Definition. A *Toda structure* on a three dimensional Einstein-Weyl space is a shear-free twist-free geodesic congruence together with a choice of homothety factor for the corresponding LeBrun-Ward gauge.

A Toda structure gives (perhaps only implicitly) a solution of the Toda equation up to changes of isothermal coordinates (x, y) and translation of z .

2.2. Proposition. *Toda structures correspond to solutions of the following closed linear system for a nonzero weight $\frac{1}{2}$ vector field \mathcal{X} and a $-\frac{1}{2}$ density σ :*

$$(2.2) \quad D\mathcal{X} = \sigma id$$

$$(2.3) \quad D\sigma = -\frac{1}{2}F^D(\mathcal{X}, \cdot) - \frac{1}{6}\text{scal}^D\langle \mathcal{X}, \cdot \rangle$$

where D is the Weyl connection, F^D is its curvature on L^1 , and scal^D is its scalar curvature, which is a section of L^{-2} . Hence Toda structures are parallel sections with respect to a natural connection on $L^{-1/2}TM \oplus L^{-1/2}$.

Proof. Equation (2.2) has already been established. Differentiating it and skew-symmetrising yields $R_{X,Y}^{D,\frac{1}{2}}\mathcal{X} = (D_X\sigma)Y - (D_Y\sigma)X$, where $R^{D,\frac{1}{2}}$ denotes the curvature of D on $L^{-1/2}TM$. Since D is Einstein-Weyl,

$$(2.4) \quad R_{X,Y}^{D,\frac{1}{2}} = -\frac{1}{6}\text{scal}^D\langle X, \cdot \rangle \Delta Y + \frac{1}{2}F^D(X, \cdot) \Delta Y - \frac{1}{2}F^D(Y, \cdot) \Delta X + \frac{1}{2}F^D(X, Y)id$$

where (for any 1-form α and vector fields X, Y) $\alpha \Delta X(Y) = \alpha(Y)X - \langle X, Y \rangle \flat \alpha$. Equation (2.3) now follows by taking a trace. \square

2.3. Corollary. *An Einstein-Weyl space admits at most a four dimensional space of Toda structures, and hence at most a three parameter family of shear-free twist-free geodesic congruences.*

By computing the curvature of the connection

$$D(\mathcal{X}, \sigma) = (D\mathcal{X} - \sigma \text{id}, D\sigma + \frac{1}{2}F^D(\mathcal{X}, \cdot) + \frac{1}{6}\text{scal}^D\langle \mathcal{X}, \cdot \rangle)$$

on $L^{-1/2}TM \oplus L^{-1/2}$, one can find obstructions to the existence of Toda structures on Einstein-Weyl spaces. In particular, substituting equation (2.3) back into $R_{X,Y}^{D,\frac{1}{2}}\mathcal{X} = (D_X\sigma)Y - (D_Y\sigma)X$, and using (2.4), yields

$$F^D(X, Y)\mathcal{X} + \langle X, \mathcal{X} \rangle F^D(Y) - \langle Y, \mathcal{X} \rangle F^D(X) = 0$$

where $F^D(X) = \flat F^D(X, \cdot)$. This condition on \mathcal{X} is simply that $\langle \mathcal{X}, \cdot \rangle \wedge F^D = 0$, or equivalently, $\langle \mathcal{X}, *F^D \rangle = 0$.

2.4. Proposition. *The congruence associated to a Toda structure on an Einstein-Weyl space (M, D) must be orthogonal to $*F^D$. Hence M admits a four dimensional space of Toda structures if and only if $F^D = 0$, i.e., the Einstein-Weyl space is Einstein.*

The sufficiency of $F^D = 0$ follows by verifying that on each of the three Einstein spaces, the Toda structures (given by Tod [13]) do indeed form a four parameter family (where one of the parameters is essentially the homothety factor of the Einstein metric).

The remaining curvature obstructions are obtained by differentiating equation (2.3) and skew-symmetrising. The resulting constraint on (\mathcal{X}, σ) is:

$$\begin{aligned} (D_X F^D)(Y, \mathcal{X}) - (D_Y F^D)(X, \mathcal{X}) + \frac{1}{3}(\langle X, \mathcal{X} \rangle D_Y \text{scal}^D - \langle Y, \mathcal{X} \rangle D_X \text{scal}^D) \\ = F^D(X, Y)\sigma. \end{aligned}$$

The Cotton-York curvature of the underlying conformal structure may be defined by $C_{X,Y}Z = (D_X F^D)(Y, Z) - (D_Y F^D)(X, Z) + \frac{1}{6}(\langle X, Z \rangle D_Y \text{scal}^D - \langle Y, Z \rangle D_X \text{scal}^D)$. Hence the above constraint relates C to $D\text{scal}^D$ and F^D . Since it is skew in X, Y , it is convenient to apply the star operator to obtain

$$\mathcal{Y}(\mathcal{X}, \cdot) + \frac{1}{6}(*D\text{scal}^D)(\mathcal{X}, \cdot) = \sigma *F^D$$

where $\mathcal{Y}(U, V) = \langle *(C_{\cdot, \cdot}U), V \rangle$ (which is well known to define a symmetric tracefree tensor). One simple consequence of this is the following refinement of Proposition 2.4.

2.5. Proposition. *The congruence associated to a Toda structure on an Einstein-Weyl space (M, D) must be orthogonal to $*F^D$ and null with respect to the Cotton-York tensor \mathcal{Y} . Hence M can only admit a Toda structure if \mathcal{Y} is indefinite on the orthogonal complement of $*F^D$.*

To see that this obstruction is nontrivial, I will apply it in the case that the Weyl structure is given by (g, ω) with ω dual to a Killing field of g . On a compact Einstein-Weyl space, there is a unique compatible metric (up to a constant) with this property, and the Einstein-Weyl structures satisfying this condition have been classified [12]. In order to avoid a case-by-case computation of \mathcal{Y} , I will derive a general formula.

2.6. Proposition. *Suppose $D = D^g + \omega$ is Einstein-Weyl with ω dual to a Killing field of g . Then*

- (i) $D_X^g F^D = \frac{1}{3} \text{scal}^D \omega \wedge \langle X, \cdot \rangle$ and so $*F^D$ is also dual to a Killing field of g .
- (ii) $\mathcal{Y}(U, V) = \frac{3}{2}(\omega(U)(*F^D)(V) + \omega(V)(*F^D)(U)) - \langle \omega, *F^D \rangle \langle U, V \rangle$.

Proof. Since $F^D = d\omega$ is closed and $D^g\omega$ is skew, $D_X^g F^D(Y, Z) = -2(R_{Y,Z}^g \omega)(X)$. The usual formulae for the Ricci tensor of g [5, 12] yield the first result by direct calculation.

Next observe that $D \text{scal}^D = D^g \text{scal}^D - 2 \text{scal}^D \omega$ and that $D_X F^D(Y, Z) = D_X^g F^D(Y, Z) - F^D(\omega \triangle X(Y), Z) - F^D(Y, \omega \triangle X(Z)) - 2\omega(X)F^D(Y, Z)$. Also, by [5], $D^g \text{scal}^D = 3D^g|\omega|^2$ and $D^g\omega = \frac{1}{2}F^D$, which leads to the following formula for C :

$$\begin{aligned} C_{X,Y} &= -\omega(X)F^D(Y, \cdot) + \omega(Y)F^D(X, \cdot) \\ &\quad + \frac{3}{2}F^D(b\omega, X)\langle Y, \cdot \rangle - \frac{3}{2}F^D(b\omega, Y)\langle X, \cdot \rangle + 2F^D(X, Y)\omega. \end{aligned}$$

Applying the star operator gives the second formula. \square

2.7. Corollary. *Suppose $D = D^g + \omega$ is Einstein-Weyl on M with ω dual to a Killing field of g . Then (M, D) cannot admit a Toda structure unless $*F^D$ is orthogonal to ω .*

Examining the explicit solutions in [12], one can easily determine for which spaces this holds: in terms of the parameters in Case 1 of [11] (which is the generic case), this condition is $abc = 0$. In particular, among the Berger spheres (given by $b = \pm c$ and $a \neq 0$), only the round sphere is Toda, verifying (in another way) the final remarks of [13].

3. TODA STRUCTURES AND SYMMETRIES

I turn now to the question: which Einstein-Weyl spaces admit more than a one dimensional family of Toda structures? In the minitwistor space picture, two Toda structures correspond to two holomorphic sections of $K_{\mathcal{S}}^{-1/2}$. Their Wronskian, being a section of $K_{\mathcal{S}}^{-1} \otimes T^*\mathcal{S} \cong T\mathcal{S}$, is a holomorphic vector field on \mathcal{S} . This symmetry of the minitwistor space induces a symmetry of the Einstein-Weyl space.

3.1. Proposition. *Suppose \mathcal{X}_1 and \mathcal{X}_2 are the weight $\frac{1}{2}$ vector fields of two Toda structures. Then $K = *(\mathcal{X}_1 \wedge \mathcal{X}_2)$ is a divergence-free twist-free conformal vector field preserving the Weyl connection.*

Proof. Differentiating K gives $DK = *(\sigma_1 \mathcal{X}_2 - \sigma_2 \mathcal{X}_1) = \frac{1}{2}d^D K$ where $D\mathcal{X}_i = \sigma_i \text{id}$. This is skew and so K is a divergence-free conformal vector field. Also $\langle K, \cdot \rangle \wedge d^D K = 0$ (since K is orthogonal to \mathcal{X}_1 and \mathcal{X}_2) so K is twist-free. Finally,

to show that K preserves the Weyl connection, it suffices to show that the Lie derivative $\mathcal{L}_K D = d \operatorname{tr} DK + F^D(K, \cdot)$ of the Weyl derivative on L^1 vanishes. Now $*F^D$ is orthogonal to \mathcal{X}_1 and \mathcal{X}_2 , so $F^D(K, \cdot) = 0$, and $\operatorname{tr} DK = 0$ since K is divergence-free. \square

Remarkably, the necessary condition of this proposition is also sufficient.

3.2. Theorem. *An Einstein-Weyl space has a two dimensional family of Toda structures if and only if it admits a (nonzero) divergence-free, twist-free conformal vector field preserving the Weyl connection.*

The necessity is the previous proposition. For the converse, suppose that K is a divergence-free, twist-free conformal vector field preserving the Weyl connection D of an Einstein-Weyl space. Since the result is local, assume K is nonvanishing. Then $DK = \alpha \triangle K$, for some 1-form α with $\alpha(K) = 0$. Furthermore, if $D^{|K|}$ is the Weyl derivative corresponding to the trivialisation of L^1 given by the length of K , then $D = D^{|K|} + \alpha$, and so $\mathcal{L}_K \alpha = 0$. Next note that, since K is twist-free, shear-free and divergence-free, it is surface-orthogonal and the integral surfaces of K^\perp are totally geodesic. The above theorem is now an immediate consequence of the following proposition.

3.3. Proposition. *Given D , K , α as above, the covariant derivative defined by $D_X^* \mathcal{X} = D_X \mathcal{X} - \alpha(\mathcal{X})X$ is flat on the bundle of vector fields of weight $\frac{1}{2}$ orthogonal to K .*

Proof. The curvature of D^* is:

$$R_{X,Y}^* \mathcal{X} = \left(-\frac{1}{6} \operatorname{scal}^D X \triangle Y + \frac{1}{2} F^D(X, \cdot) \triangle Y - \frac{1}{2} F^D(Y, \cdot) \triangle X + \frac{1}{2} F^D(X, Y) \operatorname{id} \right) (\mathcal{X}) \\ - \left((D_X \alpha)(\mathcal{X}) + \alpha(X) \alpha(\mathcal{X}) \right) Y + \left((D_Y \alpha)(\mathcal{X}) + \alpha(Y) \alpha(\mathcal{X}) \right) X.$$

Now since K is a conformal vector field preserving D , $D_X(DK) = R_{X,K}^D$. Also $DK = \alpha \triangle K$, so $D_X(DK) = (D_X \alpha + \alpha(X) \alpha) \triangle K$. Contracting with K and using the fact that $\alpha(K) = 0$ and $\mathcal{L}_K \alpha = 0$ (i.e., $(D_X \alpha)(K) = -\alpha(D_X K) = \langle K, X \rangle |\alpha|^2$) gives

$$D_X \alpha + \alpha(X) \alpha = \frac{1}{2} F^D(X) - \frac{1}{6} \operatorname{scal}^D X^\perp + |\alpha|^2 X^\parallel,$$

where X^\parallel and X^\perp denote the components of X parallel and orthogonal to K . Substituting this into the formula for R^* gives, for \mathcal{X} orthogonal to K ,

$$R_{X,Y}^* \mathcal{X} = \frac{1}{2} F^D(X, Y) \mathcal{X} - \frac{1}{2} \langle \mathcal{X}, Y \rangle \flat F^D(X, \cdot) + \frac{1}{2} \langle \mathcal{X}, X \rangle \flat F^D(Y, \cdot).$$

This vanishes for all X, Y because $F^D(K, \cdot) = 0$ and \mathcal{X} is orthogonal to K , so $\langle \mathcal{X}, \cdot \rangle \wedge F^D = 0$. \square

The parallel sections of D^* satisfy $D\mathcal{X} = \alpha(\mathcal{X}) \operatorname{id}$ and hence give a two dimensional family of Toda structures.

A consequence of this theorem is the following converse to Corollary 2.7.

3.4. Proposition. *Suppose $D = D^g + \omega$ is Einstein-Weyl on M with ω dual to a Killing field of g and that $*F^D$ is orthogonal to ω . Then $*F^D$ is dual to a*

divergence-free twist-free conformal vector field preserving the Weyl connection, and so M admits a two dimensional family of Toda structures.

Proof. Let $K = \mu_g^3 \langle *F^D, \cdot \rangle$ be the vector field dual to $*F^D$ with respect to g , where μ_g is the trivialisation of L^1 determined by g . Now by Proposition 2.6, $D^g K = -\frac{1}{3} \text{scal}^D \mu_g^3 * \omega$ (here $*\omega$ is viewed as a skew endomorphism). Since $D = D^g + \omega$, $DK = -\frac{1}{3} \text{scal}^D \mu_g^3 * \omega + \omega \Delta K + \omega(K)id$. Now if $*F^D$ is orthogonal to ω then $\omega(K) = 0$ and $-\frac{1}{3} \text{scal}^D \mu_g^3 * \omega = \alpha^g \Delta K$ for some 1-form α^g . Hence K is a divergence-free twist-free conformal vector field, and it preserves the Weyl connection since $F^D(K, \cdot) = 0$, by definition. \square

This result could also be easily established by considering each case in turn (most of which are straightforward). These spaces will feature in the final section.

4. EINSTEIN-WEYL SPACES WITH AN AXIAL SYMMETRY

In this section, I will find explicitly all the Einstein-Weyl spaces admitting a two dimensional family of Toda structures. According to the previous section, this is equivalent to classifying the Einstein-Weyl spaces admitting a divergence-free twist-free conformal vector field K preserving the Weyl connection. I will say that these spaces are Einstein-Weyl *with an axial symmetry*. On such a space, there is a two dimensional family of Toda structures given by the weight $\frac{1}{2}$ vector fields \mathcal{X} orthogonal to K and satisfying $D\mathcal{X} = \alpha(\mathcal{X})id$, where $DK = \alpha \Delta K$. In particular, $D_K \mathcal{X} = \alpha(\mathcal{X})K = D_{\mathcal{X}} K$, so $\mathcal{L}_K \mathcal{X} = 0$ and these Toda structures are K -invariant.

Pick one such Toda structure \mathcal{X} . Then $\alpha(\mathcal{X}/|\mathcal{X}|)$ is the section τ of L^{-1} given by this Toda structure, and it is only identically zero if \mathcal{X} is a parallel vector field (which can only happen on flat space). As shown by LeBrun [10], τ is a solution of the abelian monopole equation and applying the Jones and Tod construction [8] to this solution gives a hyperKähler metric with a Killing field X [1, 4]. The Toda structure is K -invariant, so K lifts to give an additional Killing field of the hyperKähler metric. Since K and X commute, some linear combination must be a triholomorphic Killing field and hence the hyperKähler metric arises via the Gibbons-Hawking Ansatz [6] from a harmonic function on \mathbb{R}^3 . This harmonic function is invariant under a Killing field of \mathbb{R}^3 and, since K is twist-free, one readily finds that this Killing field must also be twist-free (see [2]). Hence it is a rotational vector field, and the harmonic function is axially symmetric. This proves the following result.

4.1. Theorem. *Let M be Einstein-Weyl with an axial symmetry. Then if M is not flat (with translational symmetry), it is one of Ward's Einstein-Weyl spaces constructed from an axially symmetric harmonic function on \mathbb{R}^3 [16], and is therefore given explicitly by:*

$$g = (V_\rho^2 + V_\eta^2)(d\rho^2 + d\eta^2) + d\psi^2$$

$$\omega = \frac{2V_\rho V_\eta d\eta + (V_\rho^2 - V_\eta^2)d\rho}{\rho(V_\rho^2 + V_\eta^2)}$$

where $(\rho V_\rho)_\rho + \rho V_{\eta\eta} = 0$.

Note that the monopole on \mathbb{R}^3 is V_η : the choice of the integral V of V_η corresponds to the choice of the quotient of the Gibbons-Hawking metric (see [2]). This freedom involves adding multiples of $\log \rho$ to V . Note also that if $V = \log \rho$, then the monopole V_η degenerates, the Einstein-Weyl space above is \mathbb{R}^3 itself, and $\partial/\partial\psi$ is the axial symmetry.

The equation for V may be viewed as an equation on \mathcal{H}^2 , by thinking of V as being in the kernel of the conformal Laplacian on $\mathbb{R}^3 \setminus \mathbb{R}$, which is conformal to $S^1 \times \mathcal{H}^2$. More explicitly, if $v = \rho^{1/2}V$, then $v_{\rho\rho} + v_{\eta\eta} = -\frac{1}{4}\rho^{-2}v$ and so v is an eigenfunction of the Laplacian with eigenvalue $\frac{1}{8}\text{scal}_{\mathcal{H}^2}$ on the hyperbolic 2-space \mathcal{H}^2 with metric $(d\rho^2 + d\eta^2)/\rho^2$.

The original choice of Toda structure \mathcal{X} may be found by rescaling g by ρ^2 to obtain the Weyl structure in the LeBrun-Ward gauge (again, see [2]):

$$\begin{aligned} g_{LW} &= \rho^2(V_\rho^2 + V_\eta^2)(d\rho^2 + d\eta^2) + \rho^2 d\psi^2 \\ &= \rho^2(dV^2 + d\psi^2) + (\rho V_\eta d\rho - \rho V_\rho d\eta)^2 \\ \omega_{LW} &= -\frac{2V_\eta}{\rho^2(V_\rho^2 + V_\eta^2)}(\rho V_\eta d\rho - \rho V_\rho d\eta). \end{aligned}$$

The 1-form $\rho V_\eta d\rho - \rho V_\rho d\eta$ is locally exact, and may be integrated explicitly by writing $V = U_\eta$ with U axially symmetric and harmonic on \mathbb{R}^3 , so that $z = -\rho U_\rho$.

The other Toda structures come from the radial congruences on \mathbb{R}^3 centred about points on the axis of symmetry. A more democratic approach involves the relationship between these examples and Joyce's construction of torus symmetric scalar flat Kähler metrics [9] from a linear equation on hyperbolic 2-space. Indeed this linear equation, given in Proposition 3.2.1 of [9], is the Cauchy-Riemann form of the equation for V , tensored trivially with \mathbb{R}^2 . Ignoring the \mathbb{R}^2 tensor factor, simply take $x_1 = \rho$, $x_2 = \eta$, $\phi_1 = \rho V_\eta$, $\phi_2 = -\rho V_\rho$ to see that Joyce's equation is solved by axially symmetric harmonic functions. However, the advantage of his approach is that the pair $(\rho V_\eta, -\rho V_\rho)$ is identified with a section Φ of a square root of the canonical bundle of \mathcal{H}^2 satisfying an invariant equation. Now the Einstein-Weyl structure may be written

$$\begin{aligned} g &= |\Phi|^2 g_{\mathcal{H}^2} + d\psi^2 \\ \omega &= \Phi^2 / |\Phi|^2 \end{aligned}$$

and so it does not actually depend upon the choice of coordinates (ρ, η) identifying \mathcal{H}^2 with the upper half plane. Such an identification is given by a choice of a point at infinity on the hyperbolic disc and each point in this circle gives a Toda congruence.

Two solutions of Joyce's equation generate a scalar-flat Kähler metric with two Killing fields, and Ward's Einstein-Weyl spaces arise as the quotients by each of these Killing fields. Joyce finds the solution $V = \log \rho$ (which generates \mathbb{R}^3) and superposes it with its image under isometries of hyperbolic 2-space (where these

isometries are applied to Φ). In this way he obtains torus symmetric selfdual conformal structures on $k\mathbb{C}P^2$, generalising (for $k \geq 4$) the torus symmetric examples obtained from the hyperbolic Ansatz of LeBrun [10].

5. EXAMPLES

The simplest axially symmetric harmonic functions on \mathbb{R}^3 are the constant functions and the fundamental solutions. The most trivial solution $V_\eta = 0$, $V = \log \rho$ yields \mathbb{R}^3 . If $V_\eta = b$ or $V_\eta = c/\sqrt{\rho^2 + \eta^2}$ then the Gibbons-Hawking metric is \mathbb{R}^4 and the triholomorphic Killing field is an infinitesimal translation or selfdual rotation respectively. Hence the Einstein-Weyl spaces obtained are the quotients of \mathbb{R}^4 by Killing fields (infinitesimal translations or rotations) given in [11].

To obtain more complicated examples, one can take linear combinations of fundamental solutions and constant solutions. In this way, one can find the Einstein-Weyl quotients of the Taub-NUT and Eguchi-Hanson metrics, more or less by direct substitution, although more manageable expressions are obtained after transforming the (ρ, η) coordinates.

The Taub-NUT solutions are given by $V = a \log \rho + b\eta + c \log \frac{\eta + \sqrt{\rho^2 + \eta^2}}{\rho}$ and it is convenient to set $\rho = r \cos \theta$, $\eta = r \sin \theta$ so that $\rho V_\eta = (br + c) \cos \theta$ and $\rho V_\rho = a - c \sin \theta$. Then

$$g_{LW} = ((br + c)^2 \cos^2 \theta + (a - c \sin \theta)^2)(dr^2 + r^2 d\theta^2) + r^2 \cos^2 \theta d\psi^2$$

$$\omega_{LW} = -\frac{2(br + c)}{r((br + c)^2 \cos^2 \theta + (a - c \sin \theta)^2)} d(-ar \sin \theta + \frac{1}{2}br^2 \cos^2 \theta + cr).$$

Note that $bc = 0$ gives the quotients of \mathbb{R}^4 mentioned briefly above.

The Eguchi-Hanson solutions are obtained from

$$V = a \log \rho + \frac{1}{2}(b + c/\varepsilon) \log \frac{\eta - \varepsilon + \sqrt{\rho^2 + (\eta - \varepsilon)^2}}{\rho} + \frac{1}{2}(b - c/\varepsilon) \log \frac{\eta + \varepsilon + \sqrt{\rho^2 + (\eta + \varepsilon)^2}}{\rho},$$

where $\varepsilon^2 = \pm 1$ (without loss of generality). When $\varepsilon^2 = -1$ this is the potential for an axially symmetric circle of charge, while $\varepsilon^2 = +1$ corresponds to two point sources on the axis of symmetry. These cases are sometimes referred to as Eguchi-Hanson I and II respectively. The former is always incomplete, but its Einstein-Weyl quotients are perhaps more interesting than those of Eguchi-Hanson II.

Coordinates adapted to these geometries are obtained via $\rho = \sqrt{R^2 - \varepsilon^2} \sin \theta$ and $\eta = R \cos \theta$ so that

$$\rho V_\eta = \frac{(bR + c \cos \theta) \sqrt{R^2 - \varepsilon^2} \sin \theta}{R^2 - \varepsilon^2 \cos^2 \theta}$$

and

$$\rho V_\rho = \frac{a(R^2 - \varepsilon^2 \cos^2 \theta) - b(R^2 - \varepsilon^2) \cos \theta + cR \sin^2 \theta}{R^2 - \varepsilon^2 \cos^2 \theta}.$$

The Toda structure is now given by:

$$g_{LW} = ((a \cos \theta - b)^2(R^2 - \varepsilon^2) + (aR + c)^2 \sin^2 \theta) \left(\frac{dR^2}{R^2 - \varepsilon^2} + d\theta^2 \right) \\ + (R^2 - \varepsilon^2) \sin^2 \theta d\psi^2 \\ \omega_{LW} = -\frac{2(bR + c \cos \theta)}{(a \cos \theta - b)^2(R^2 - \varepsilon^2) + (aR + c)^2 \sin^2 \theta} d(-aR \cos \theta + bR - c \cos \theta).$$

The family given by $a = 0, \varepsilon^2 = -1$ also arises as a quotient of the scalar flat Kähler metric on $S^2 \times \mathcal{H}^2$. If we write this as:

$$g = \frac{dR^2}{R^2 + 1} + (R^2 + 1)ds^2 + d\theta^2 + \sin^2 \theta d\phi^2$$

then $K = b\partial/\partial s + c\partial/\partial \phi$ is a Killing field. Coordinates adapted to K are given by $\chi = bs + c\phi, \psi = b\phi - cs$ so that K is a multiple of $\partial/\partial \chi$ and the quotient metric $g - g(K, \cdot)/g(K, K)$ is

$$\frac{dR^2}{R^2 + 1} + d\theta^2 + \frac{(R^2 + 1) \sin^2 \theta}{b^2(R^2 + 1) + c^2 \sin^2 \theta} d\psi^2.$$

This is the same conformal structure as above, and one readily checks that the Weyl structures also agree. Now $S^2 \times \mathcal{H}^2$ is conformal to $\mathbb{R}^4 \setminus \mathbb{R}$ and so these Weyl structures are globally defined on S^3 for $b \neq 0$ (since $\partial/\partial s$ is a dilation). Hence, as remarked in [2], these quotients of \mathbb{R}^4 by dilation plus planar rotation are Toda (although the congruences are not globally defined on S^3). Additionally, the calculations of this section verify explicitly that they are Einstein-Weyl with an axial symmetry, in accordance with Proposition 3.4, and arise from the Eguchi-Hanson I metrics.

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